1. Large Deviations and Population Dynamics

Unusual time averages

We ask for the probability that it sustains an unusual time-average for an observable during a long interval:
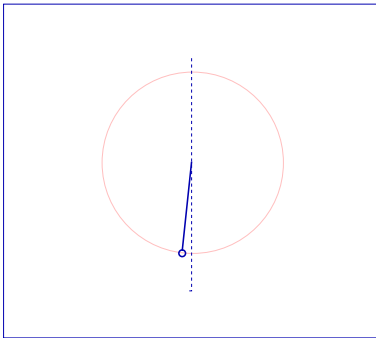
$$\frac{1}{t} \int dt' \, A(t') = \overline{A}$$

The most celebrated example is the average power, which can be interpreted in some cases as entropy production:
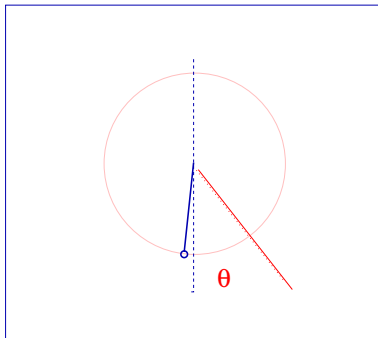
$$\sigma_t = -\frac{1}{t} \int_0^t dt' \, \frac{\mathbf{f}\dot{\mathbf{q}}}{T}$$

$$P(A) = \sum_{\text{Trajec.}} (\text{Prob. Trajectory}) \, \delta \left(t\overline{A} - \int_o^t dt' \, A(t') \right)$$
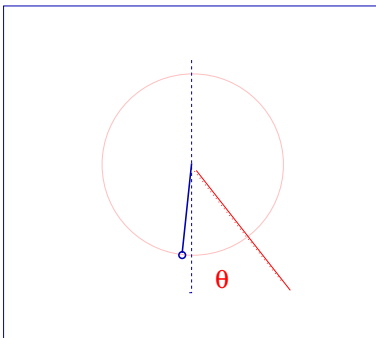
a pendulum immersed in a low-temperature bath
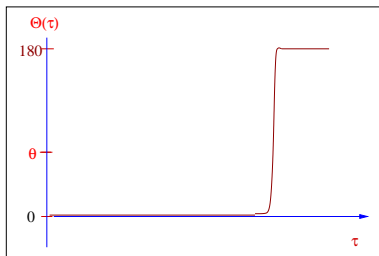
a pendulum immersed in a low-temperature bath

and ask: *given* **that during a specific minute the average angle took the value** θ

i) **what is the probability of this?**   *ii)* **how does this come about?**

Imposing the average angle, the trajectory shares its time between saddles 0° and 180°

phase-separation, a first order transition!

Consider a process:

$$\frac{dP(q,t)}{dt} = -LP(q,t)$$

$$P(\mathbf{q_o}, t_o \to \mathbf{q}, t) \quad = e^{-Lt}P(q,0) = \int D[q] \quad (\text{Prob. Trajectory})$$

$$= \int D[q] \; e^{-\mathcal{L}_t(trajectory)}$$

Unusual time averages

We ask for the probability that it sustains an unusual time-average for an observable during a long interval:

$$\frac{1}{t} \int dt' \, A(t') = \overline{A}$$

The most celebrated example is the average power, which can be interpreted in some cases as entropy production:

$$\sigma_t = -\frac{1}{t} \int_0^t dt' \, \frac{\mathbf{f}\dot{\mathbf{q}}}{T}$$

$$P(A) = \sum_{\text{Trajec.}} (\text{Prob. Trajectory}) \, \delta \left(t\overline{A} - \int_o^t dt' \, A(t') \right)$$

Writing the delta function as an exponential

$$\delta\left(t\overline{A} - \int_o^t dt' \; A(t')\right) = \int_{-i\infty}^{+i\infty} d\mu \; e^{\mu\left(t\overline{A} - \int_o^t dt' \; A(t')\right)}$$

$$
\begin{aligned}
P(A) &= \int_{-i\infty}^{+i\infty} d\mu \; e^{\mu t \overline{A}} \quad \underbrace{\int D[\mathbf{q}] \; (\text{Prob. Trajectory}) \, e^{-\mu \int_o^t dt' \; A(t')}}_{\displaystyle \downarrow} \\
&= \int_{-i\infty}^{+i\infty} d\mu \; e^{\mu t \overline{A}} \qquad\qquad\quad \times \; e^{-t G(\mu)}
\end{aligned}
$$

which defines the large-deviation function $G(\mu)$.

For example, in the Fokker-Planck case, it reads:

$$e^{-tG(\mu)} = \int D[\mathbf{q}] \ e^{\int dt' \ [\sum_i - \mathcal{L} - \mu \int_o^t \ dt' \ A(t')]}$$

What we have done is nothing but the analogue of a passage from a microcanonical calculation of 'entropy' $= \ln \bar{A}$, to a canonical calculation of 'free energy' $G(\mu)/\mu$ at 'inverse temperature' μ. The 'space' in our problem is in fact the time, and 'extensive quantities' are those that are proportional to time: we extracted a time in the definition of $G(\mu)$ in order to make it 'intensive'.

For large t, in analogy with the thermodynamic limit, assuming that $G(\mu)$ has a good limit, we may evaluate the integral over μ by saddle point, to obtain:

$$\ln P(A) \sim t[\mu^* A - G(\mu^*)]$$

$$A(\mu^*) = \left. \frac{dG}{d\mu} \right|_{\mu^*}$$

This is the Legendre transform taking from canonical to microcanonical.
Now, by simple comparison, operator language:

$$e^{-tG(\mu)} = \langle final|e^{-t[L+\mu A(\mathbf{q})]}|init\rangle$$

Note the fundamental difference between this large-deviation functions and those of Friedlin-Wentzell: in that case by 'large'

deviations we meant that they are exponentially small in the temperature (or the coarse-graining size) while here we mean that

they are sustained for long times, and the only large parameter is precisely the time t.

Simulating large deviations – relation with selection

$$\dot{P} = -[L + \mu A]P \quad \to P(\mathbf{q}, t) = e^{-t[L+\mu A]}$$

We are dealing with a dynamics without probability conservation. In fact, we can reproduce it by using a large number of non-interacting walkers, each performing the original (Langevin) dynamics with independent noises, occasionally giving birth to another walker starting in the same place, or dying. A negative (positive) value of $\mu A(\mathbf{q})$ gives a probability $|\mu A(\mathbf{q})|dt$ of making a clone or of dying, respectively, in a time-interval dt. At each time, the global number of clones $M(t)$ changes, in such a way that for long times $M(t)/M(t=0) \sim e^{-\lambda_{min}(\mu)t}$. In practice, one can normalise the total number periodically by cloning or decimating all walkers with a random factor. The factor needed to keep the population constant is, again, the exponential of the lowest eigenvalue.

N   independent   simulations

with probability c . A per unit time kill or clone

...   continue ...
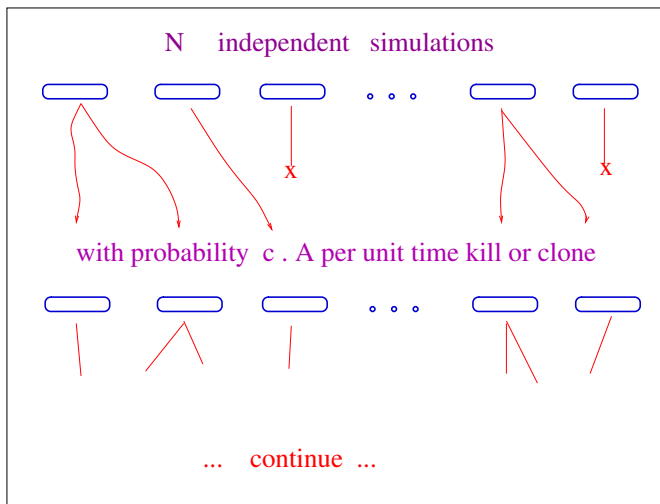
a way to count trajectories weighted with e^{cA}
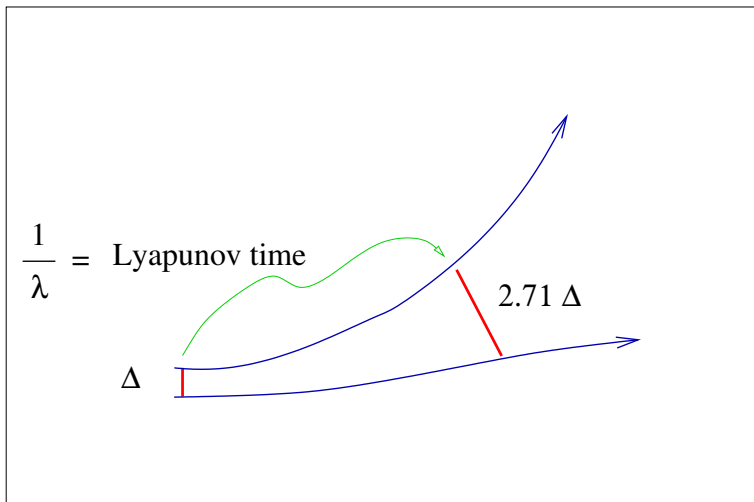
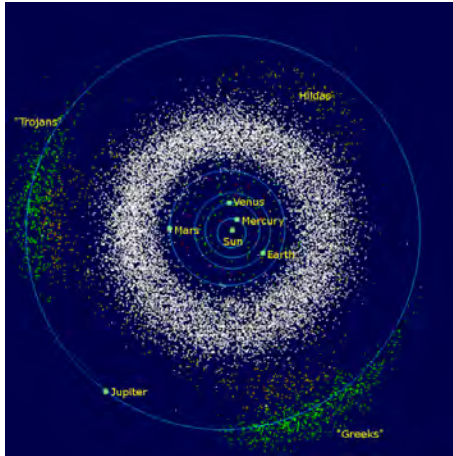**Trajectories with rare values of Lyapunov exponents: vector process**

à la Diffusion Monte Carlo: particles have a vector $\vec{u}$ attached, which evolves with

$$\dot{u}_i = -\frac{\partial^2 V}{\partial x_i x_j} u_j$$

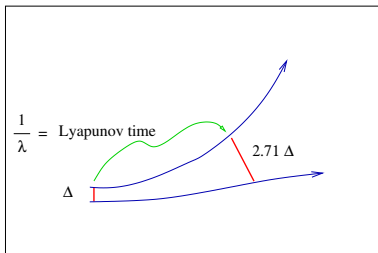**Particles perform ordinary diffusion + cloning $\propto \frac{d|u|}{dt}$**
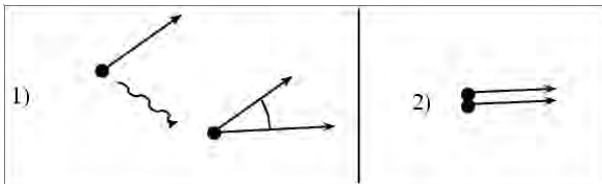
2. Large Deviations:

an Advertisement

because planets disturb one another, the dynamics is chaotic

chaotic?

difference between trajectories multiplies by $e = 2.71...$ **every** $\sim 5M$ **years** Laskar

$\lambda \equiv 1/5MYrs$ **is called the** *Lyapunov exponent*

$\lambda > 0 \;\; \rightarrow$ **chaos**

Poincaré section: a better visualisation:
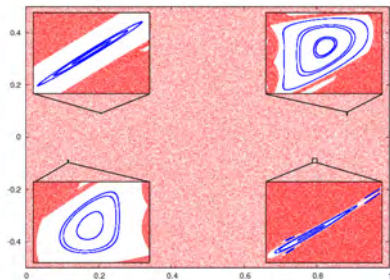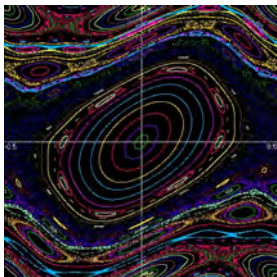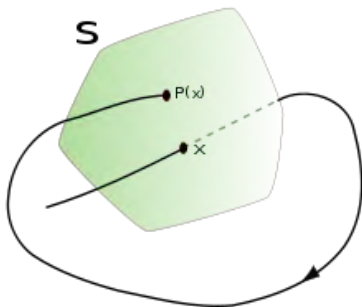
Consider this:

- The solar system formed $\sim$ 4.5 GYr ago
- starting from the present conditions, *depending on details*, 1% of histories run into trouble in 5 GYr <small>Laskar and Gastineau</small>
- If you start a random planetary system in your computer, almost always it quickly runs into trouble.
- If you observe a planetary system, many conditions within the observational error imply recent formation or immediate destruction

You need to know rare trajectories
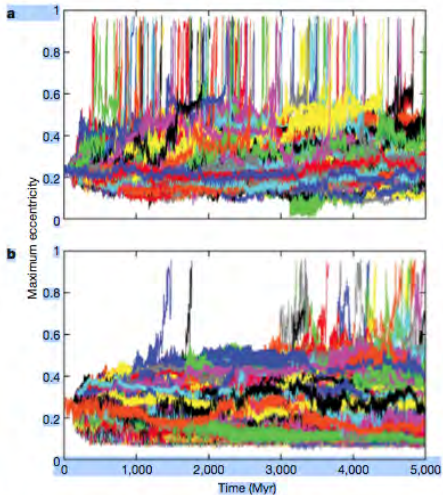
Figure 1 | Mercury's eccentricity over 5 Gyr. Evolution of the maximum

Laskar et al

Another example

Draupner wave record
January 1 1995 at 15:20, hs = 11.3m
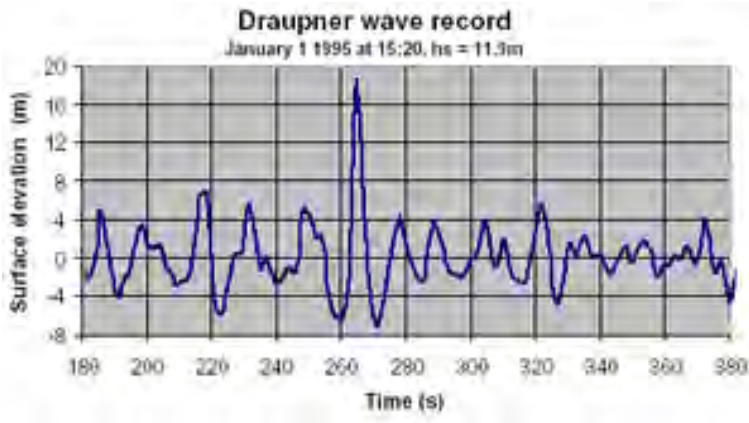
The Draupner rogue wave Taylor, Wiki

Another example

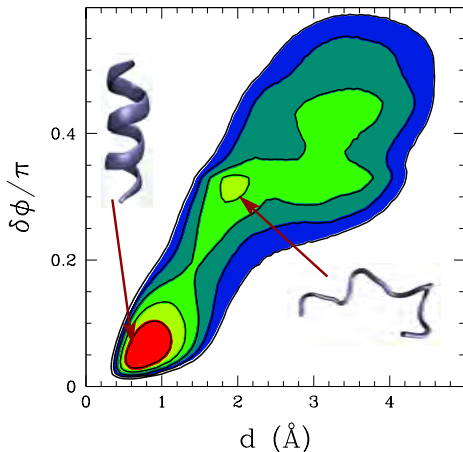**The Draupner rogue wave** Taylor, Wiki

Another example

T. Duriez, J.-L. Aider, E. Masson, J.E. Wesfreid ; Qualitative investigation of the main flow features over a TGV ; Proceedings of the Euromech Colloquium 509, Vehicle Aerodynamics, Berlin, Allemagne, **2009**, p. 52-57
http://opus.kobv.de/tuberlin/volltexte/2009/2249/
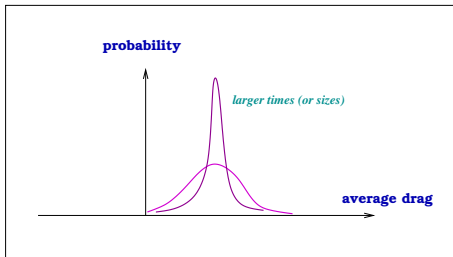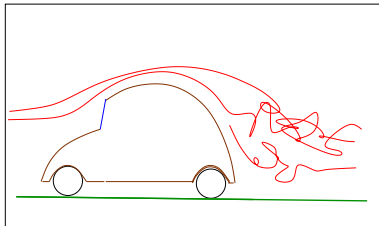
$$\bar{f}_\tau = \frac{1}{\tau} \int_0^\tau f(t)\, dt$$

$$\bar{f}_{\tau} = \frac{1}{\tau} \int_0^{\tau} f(t) \ dt$$

peptide helix-coil transition most trajectories spend their time around one configuration

We wish to simulate an event with an unusually large value of A

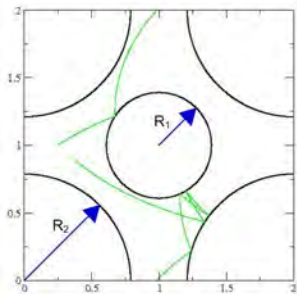**without having to wait for this to happen spontaneously**

but without forcing the situation artificially

N   independent   simulations

x   x

with probability c . A per unit time kill or clone

...   continue ...
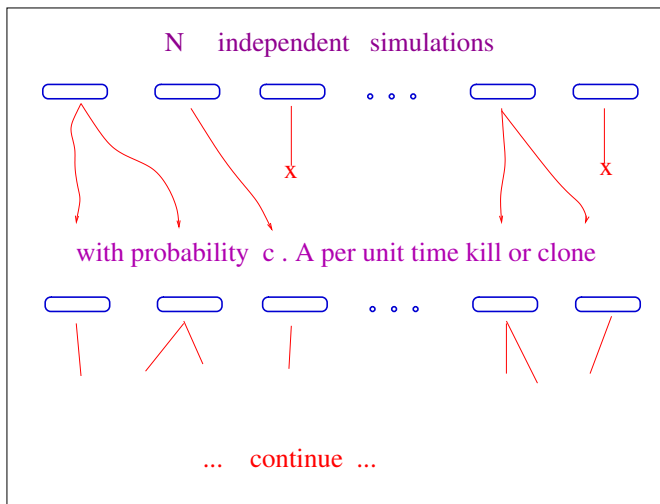
a way to count trajectories weighted with e^{cA}

Some examples

Driven Lorentz Gas
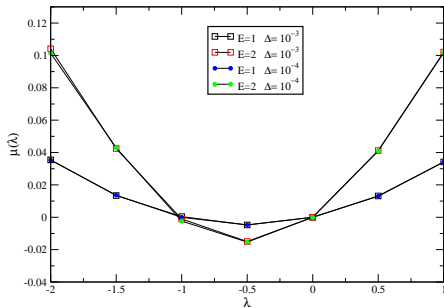
Fermi-Pasta-Ulam chain

$$H = \sum_{i=1}^{N} \left(\frac{1}{2} p_i^2 + \frac{1}{2}(x_i - x_{i+1})^2 + \frac{\beta}{4}(x_i - x_{i+1})^4 \right)$$
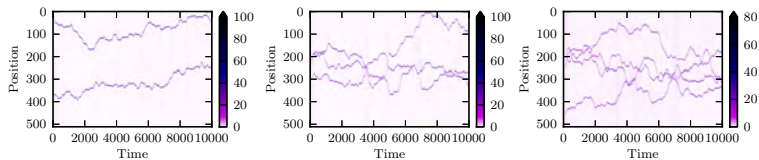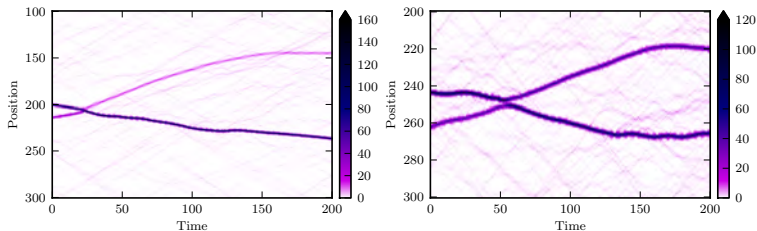
$$A = t * \lambda$$

typical
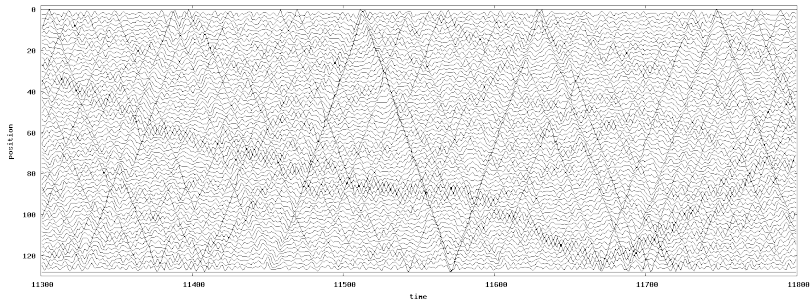
unusually unchaotic
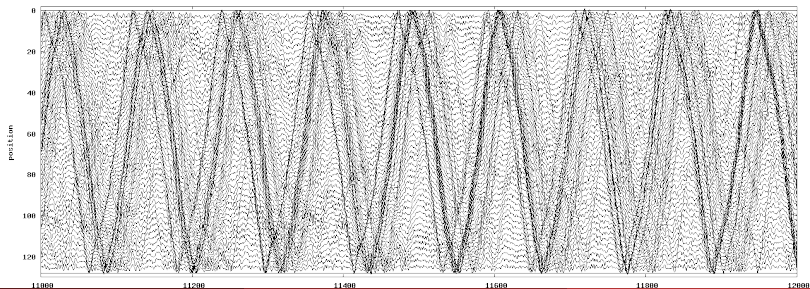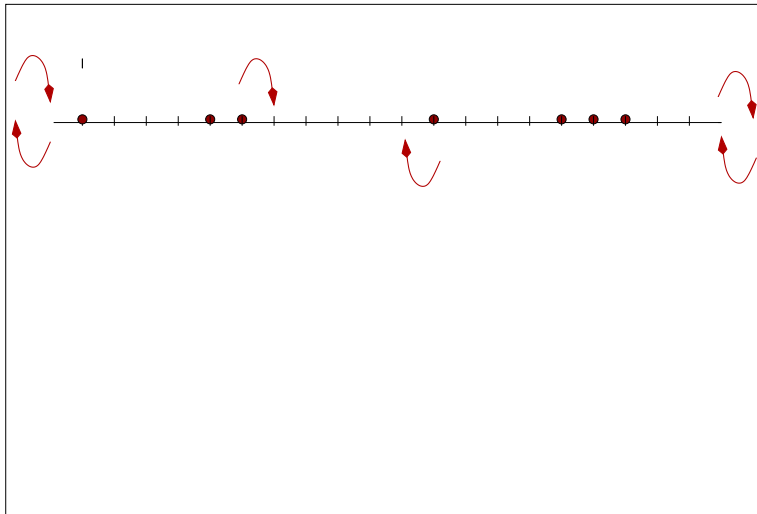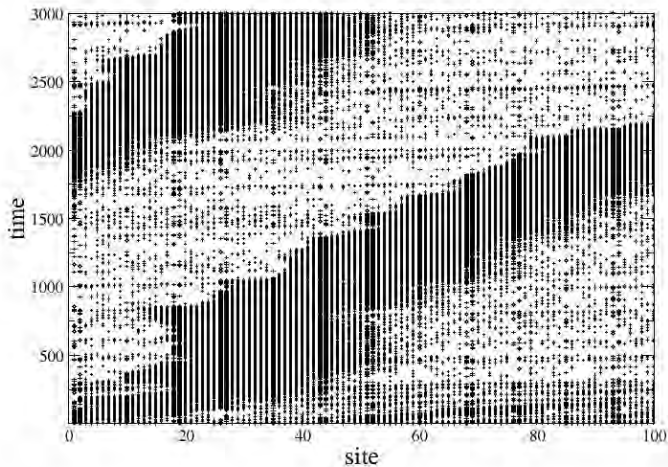
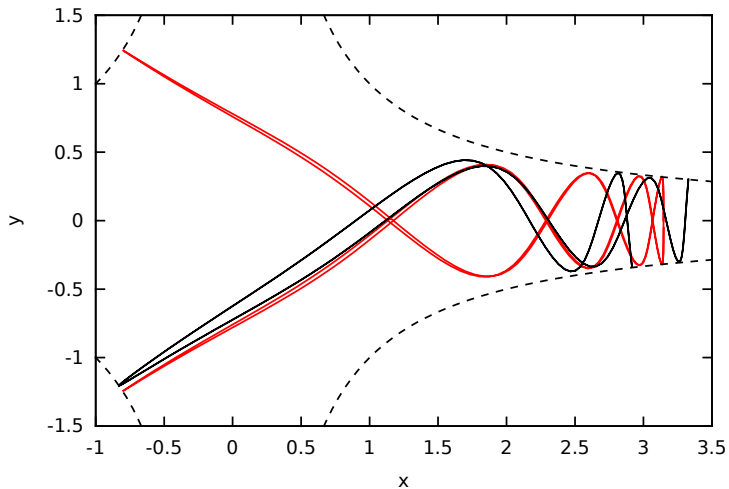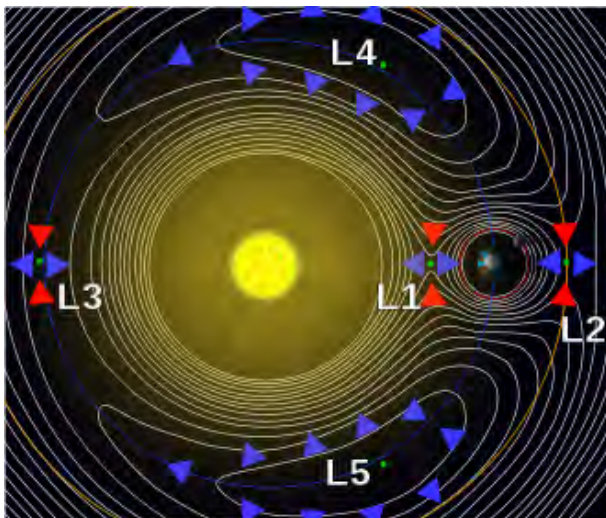unusually chaotic: breathers (Tailleur et al)

The Symmetric Exclusion Process

The SEP: unusually low current (a jam)

The x^2y^2 potential: regular islands

Lagrange Point

Effect of eccentricity and mass ratio